Terrestrial Planets

Week 8

Professor Olivia Jensen Earth and Planetary Sciences FD Adams 131C



Orbits

- During the next two weeks, we will learn something about the orbits of planets and moons and how we may use our knowledge of these orbits to infer their masses and the mass of the Sun.
- Note that <u>Vera Rubin</u>, she who confirmed the existence of dark matter and who might be seen as the "star" of the video <u>Most of the Universe is</u> <u>Missing</u>, died December 25, 2016.



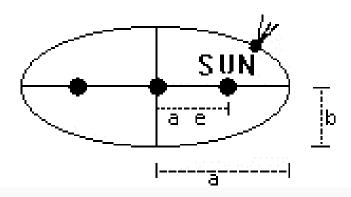
Orbital dynamics, Kepler's Laws and Newtonian Gravity

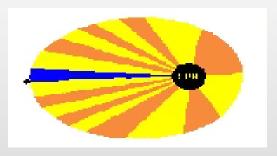
While the ancient Greeks knew that the Earth was spherical – they actually had a very good idea of its size too – and that the planetary system, then known to comprise only the 6 inner planets, was most simply viewed from a heliocentric (Sun-centred) perspective, Europeans did not accept this knowledge until Nicolas Copernicus in the early 1500s and Galileo Galilei a century later promoted the heliocentric viewpoint*.

*Practically and more accurately, it would be well to consider the centre of mass of the Solar System of Sun and planets to be fixed and the Sun and all the planets in orbit about this "*barycentre*".

Kepler's Laws

Kepler's first law: The orbit of a planet about the Sun is an ellipse with the Sun at one focus.





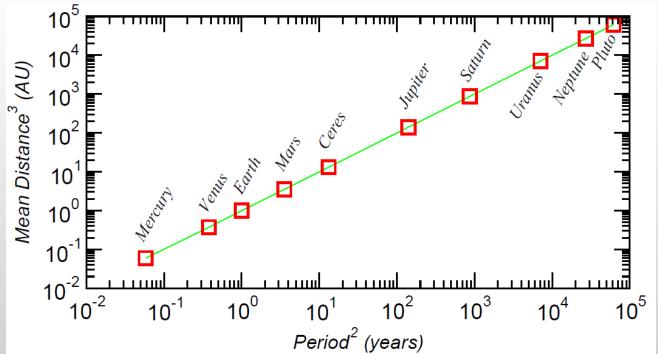
Kepler's second law: A line joining a planet and the Sun sweeps out equal areas in equal intervals of time.



Animations: Bill Drennon, Central Valley Christian High School, Visalia, CA

Kepler's Laws

Kepler's third law: The square of a planet's siderial period (i.e., referenced to the distant stars) is proportional to the cube of the length of its orbit's semi-major axis. That is $T^2 \propto a^3$ where T is the orbital period and a is the semi-major axis of the orbital ellipse.





Kepler's Laws



https://www.youtube.com/watch?v=P7vc4e8efus



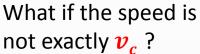
Newton's Gravity

Kepler's Laws were studied by <u>Isaac Newton</u> (1642-1727). He recognized that a field of force which he called *gravity* was fundamentally responsible for the ordered, Keplerian, motion of the planets. In his *Philosophae naturalis principia mathematica*, published in 1687, he showed how the force of gravity constrained planets into precisely Keplerian orbits.



Orbits and Gravity

What speed v_c allows for a stable circular orbit? v_c m_{p} $M_{\odot}m_p$ *Gravity* provides $F_g = G$ the *centripetal* force that holds the planet in M_{\odot} circular orbit. F_{c} $F_c = F_g$



 $v < v_c$: in spiral $v = 2v_c$: parabola $v_c < v < 2v_c$: ellipse $v > 2v_c$: escape hyperbola



Weighing the Sun

Gravity balances the necessary centripetal force holding a planet in orbit:

$$G \frac{M_{\odot}m_p}{r^2} = \frac{m_p v_p^{-2}}{r}$$
Noting that for Earth, the orbital speed is
$$v_p = \frac{2\pi r}{T},$$
where *T* is our 1-year orbital period.
With a little mathematical reduction, we find that
$$T^2 = \frac{4\pi^2 r^3}{GM_{\odot}},$$
a statement of Kepler's 3rd Law. Knowing *G*, , measuring *T* and
r, we find the mass of the Sun:
$$M_{\odot} = \frac{4\pi^2 r^3}{GT^2}$$

Weighing the Earth

Gravity provides the necessary centripetal force holding the Moon in orbit about the Earth. We only need know the period of the Moon's orbit about the Earth, or more properly about the Earth-Moon **barycentre*** in order to determine Earth's mass.

The Moon orbits the Earth in **1 month** (by definition). **But**, what is the length of the month? There are several measures of the month but only one of them is the appropriate measure of the time of orbit.

*The <u>Earth-Moon barycentre</u> is displaced from the centre of mass of Earth. For *r*, the Earth-Moon distance, the displacement of the barycentre from the Earth's centre of mass in the direction toward the Moon is obtained as:

$$d=rac{M_c}{M_{\oplus}}rpprox$$
 4636km

Barycentre of the Solar System? All the planets tug on the Sun and as they orbit, the Sun too orbits the barycentre of the Solar System.



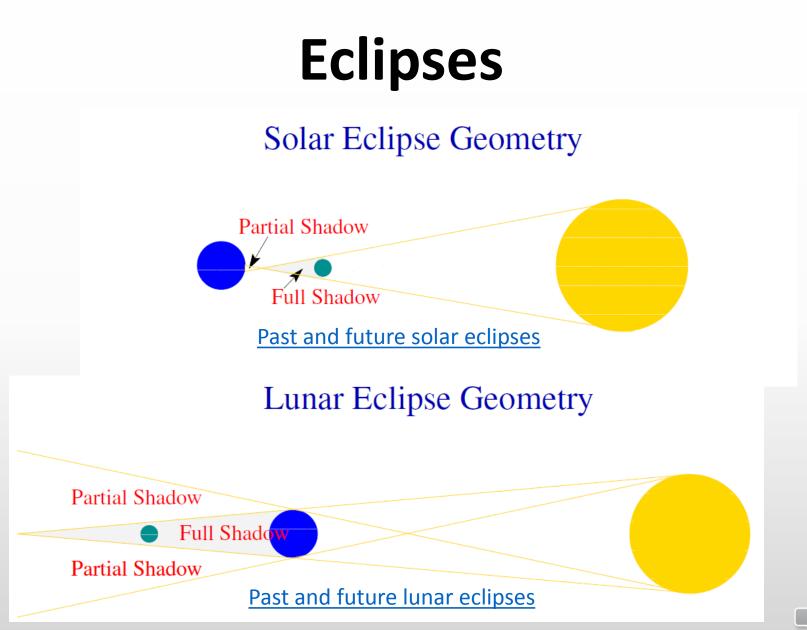
Measure of the Month

Lunar (or Synodic) Month: Once each lunar month, the Moon comes to Full Moon phase. This happens during each month at the moment when the line from the Sun-to-Earth-to-Moon is most nearly a straight line. (Mean period: 29.531 days)

<u>Siderial month</u>: The interval between two successive times at which the Moon appeared in the same location relative to the distant stars in the night sky. (Mean period: **27.322 days**)

Draconitic month: The interval between two successive times that the Moon appears highest in the night sky. This is referenced to Earth geography, to the latitude of the observer. (Mean period: **27.212 days**)

<u>Tropic month</u>: Referenced to the Earth's rotation axis that appears to pierce the celestial sphere near Polaris, the North Star, we measure the angular distance between the Moon and this axis. The angular distance repeats in 1 Tropic Month. (Period: **27.331(582) days**) *The tropic month is the orbital period of the Moon*.





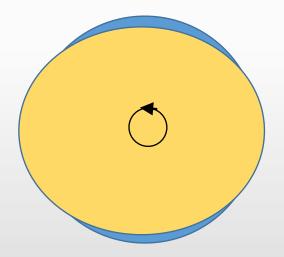
Tidal forces on Earth

The Moon and Sun exert differential tidal forces on the body, water and atmosphere of Earth. On that side of the Earth closest to the Moon or Sun, the gravitational pull towards the Moon or Sun is greater than it is on the opposite. This creates a differential force across the body of Earth and the body of the Earth, its oceans and atmosphere adjust through what we observe as "*tides*".

There are **body**, **ocean** and **atmospheric tides**. Ocean and atmospheric tides involve movable fluids and are, thus, more difficult to calculate than is Earth's body tide.

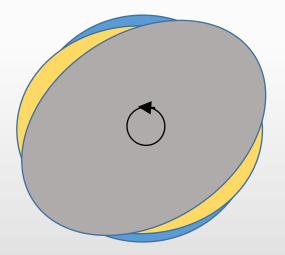
Let us look into the body tide.





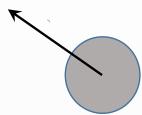
The solar bulge is always directed towards the Sun (noon) or nadir (midnight). It follows a 24 hour cycle.

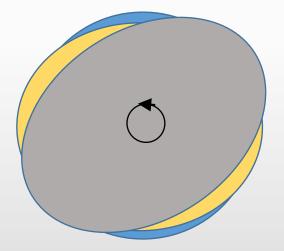




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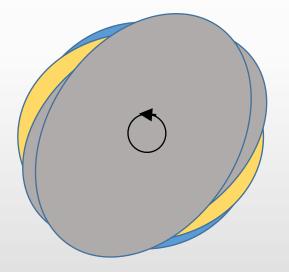
The **lunar tidal bulge** is aligned with the Earth-Moon line.





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Due to the fact that the Moon is in orbit about the Earth, the **lunar tidal bulge** which follows the Earth-Moon line is delayed relative to the solar bulge by an average of 51 minutes per day.

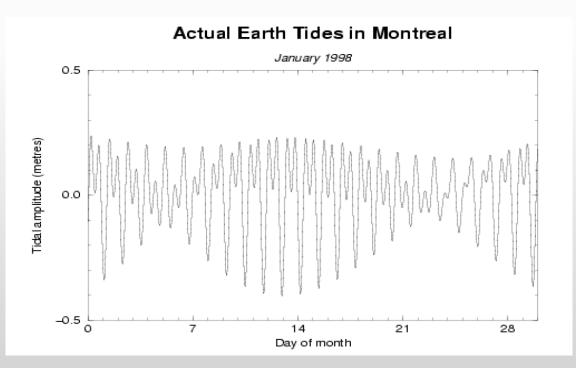


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The **lunar tidal bulge** is delayed by an average of 51 minutes per day. The tide we feel is the sum of the two effects at any time.

Tidal beating

The 24-hour Solar tide, with two high tides and 2 low tides each day, beats with the 24 hour 51 minute lunar tide. The lunar tide is not exactly periodic at 24 hour 51 minute because the Lunar orbit is inclined relative to Earth's spin axis. The two tidal forcings add together to produce the measureable body tide.





Black hole at Galactic Centre

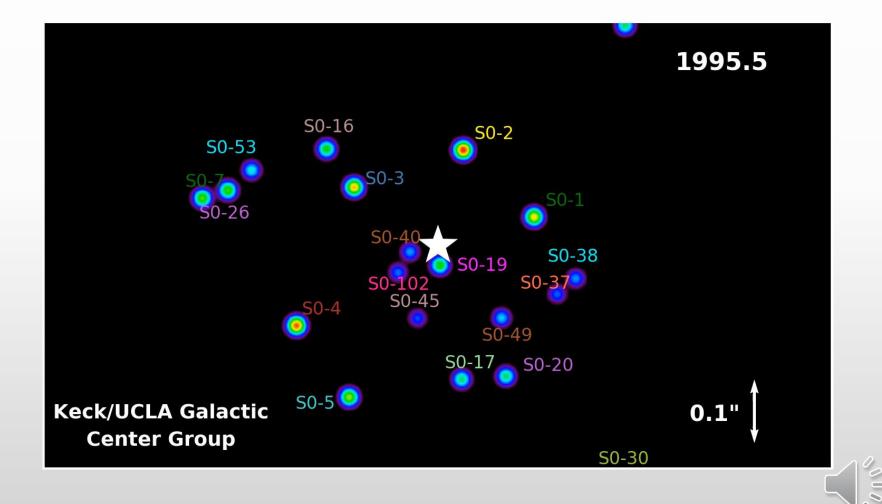
<u>Andrea Ghez</u>, Nobel Prize (2020) of UCLA, used a procedure, very similar to that described previously for "weighing" the Sun, to measure the unseen mass at our galactic centre.

Having observed the orbits and periods of stars in the direction of the galactic centre, she and her colleagues determine that there is an "invisible" supermassive gravitational centre, having a mass of almost $4 \times 10^6 M_{\odot}$, at the centre of our Milky Way.

Andrea Ghez' exploration of the central galactic black hole



Black hole at Galactic Centre



The Mass of Dark Matter

<u>Vera Rubin</u> looked at the orbital speeds of stars well outside the core zone of many galaxies. She studied the "*rotation curve*" to show that stars at distance from galactic centres were orbiting *much faster* than could be accounted for by the observable baryonic matter of the galaxies.

Our analysis above obtains the *expected velocity* of orbit due to mass within the orbital path:

$$v_T = \sqrt{\frac{GM_c}{r}}$$

She confirmed the existence of "*Dark Matter*" which seems to comprise about 28% of everything in the Universe.

Most of the Universe is Missing?



Online verison here: https://www.dailymotion.com/video/x226e38

